

# Remote Sensing & Photogrammetry W4

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# Image processing

## 1. Visual interpretation of single spectral band

- Readout of DN and coordinates: x,y

## 2. Image enhancement

- Histogram calculation
- Linear histogram stretching
- Image comparison before and after *stretching*
- Different parameters of linear *stretching*
- Histogram saturation and equalization

## 3. Visualization of multispectral bands

- RGB
- FCC

# Image processing

## 4. Multi bands operation

- Ratio – Vegetation Index (VI)
- Normalized ratio – Normalized Vegetation Index (NDVI)
- Multi channel statistics
- Principal Component Analysis (PCA)
- Map algebra
- Image fusion

## 5. Image classification

- Density slicing
- piece-wise linear stretching
- Multispectral image classification
  - Sampling
  - Different algorithms
  - Classification accuracy assessment
  - Post classification operation

# Map algebra - radiance

- Radiance (luminance) calculation:

$$L_{\lambda} = gain \cdot DN + offset = ((L_{max} - L_{min}) / 255) \cdot DN + L_{min}$$

- $L_{\lambda}$  – spectral radiance recorded by sensor
- $L_{min}$  – minimal radiance of detectors in PAN: – 5.00
- $L_{max}$  – maximal radiance of detectors in PAN: 244.00
- DN – of channel 8, PAN

# Map algebra - albedo

Albedo

$$\rho = \frac{\pi \cdot L_{\lambda} \cdot d^2}{ESUN_{\lambda} \cdot \cos \theta_S}$$

where:

$\rho$  – albedo

$L_{\lambda}$  – spectral radiance recorded by sensor

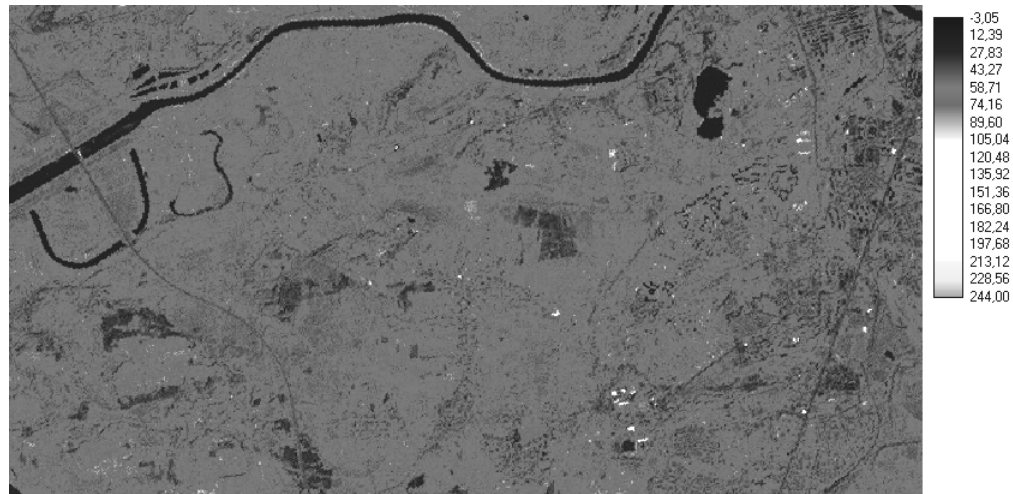
$d$  – distance between Earth and Sun in astronomical units  
for given day of the year

$ESUN_{\lambda}$  – mean irradiance = 1368.00

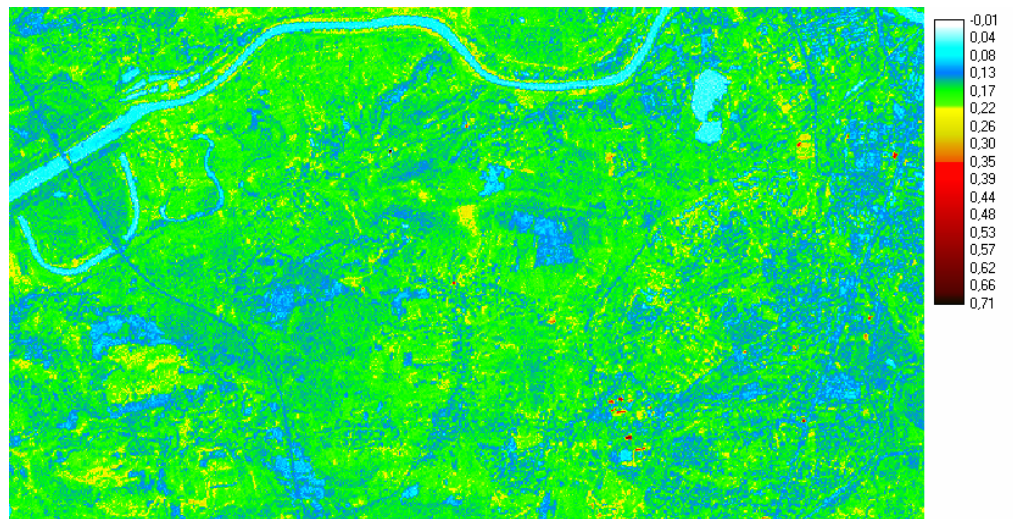
$\theta_S$  – Sun zenith angle

# Map algebra - albedo

DN  
PAN  
Band 8  
Landsat ETM+



Albedo  
(0.52 – 0.90  $\mu\text{m}$ )



# VIR, SWIR, TIR

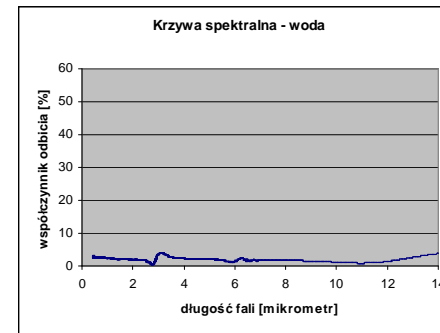
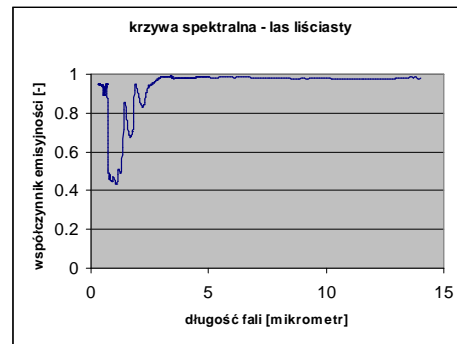
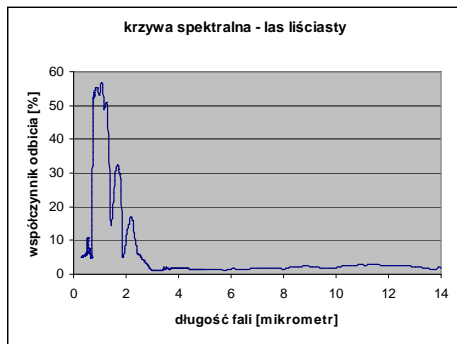
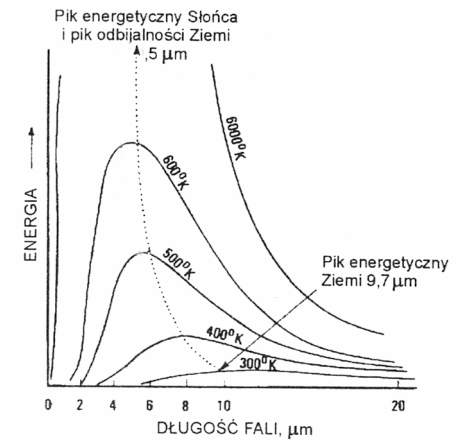
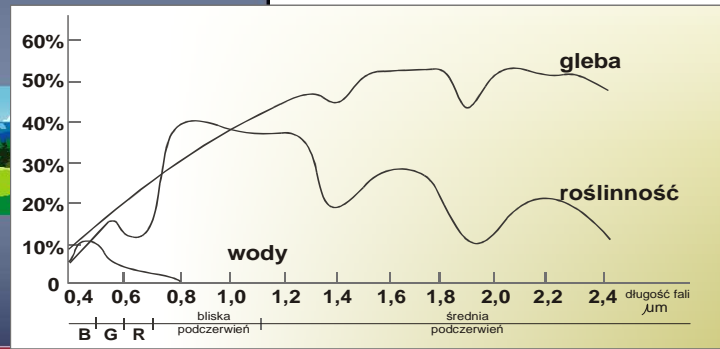
- Współczynnik odbicia
- Współczynnik absorpcji (emisyjności)
- Współczynnik transmisji

$E = E_{\rho} + E_{\tau} + E_{\alpha}$   
 $\rho = E_{\rho}/E$   
 $\tau = E_{\tau}/E$   
 $\alpha = E_{\alpha}/E$   
 $1 = \rho + \tau + \alpha$

Prawo Kirchhoffa:  $\alpha = \epsilon$      $\rho = 1 - \alpha$

**$\rho = 1 - \epsilon$**

## Spectral curves



$$E_{\rho}(\lambda) = \rho(\lambda) E(\lambda)$$

$$E_s(\lambda) = \epsilon(\lambda) B(\lambda, T_s) + (1 - \epsilon(\lambda)) E(\lambda)$$

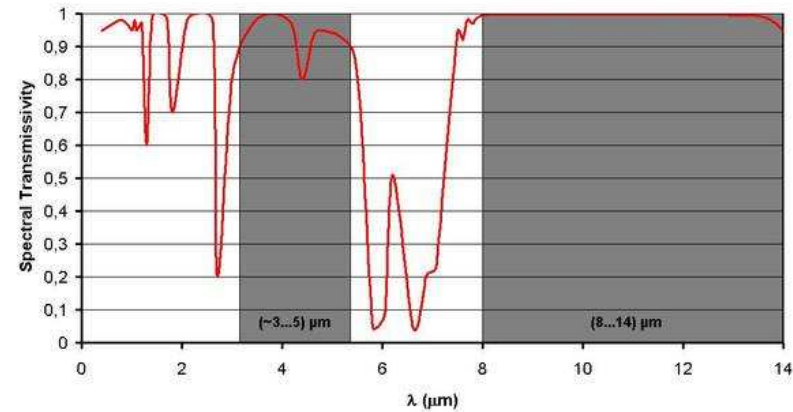
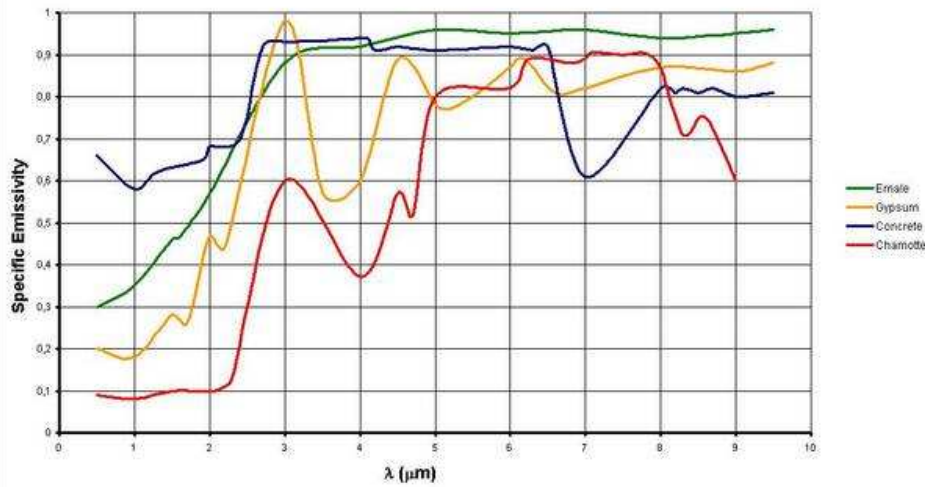
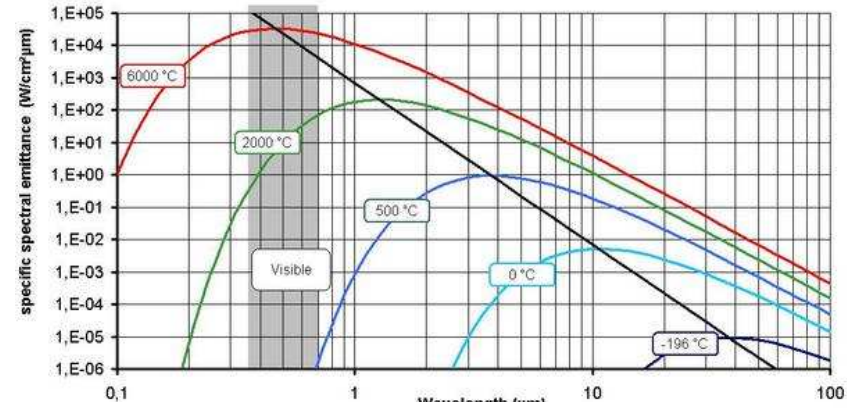
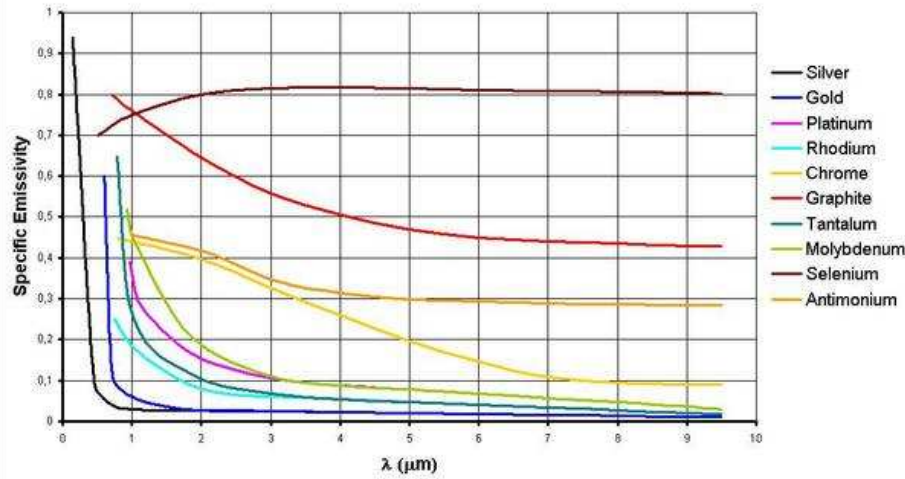
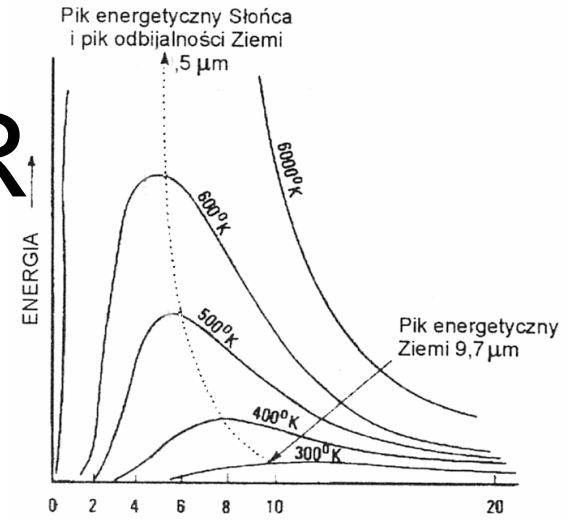
# EM laws TIR

$$M_\lambda = \frac{c_1}{\lambda^5 * [\exp(c_2/\lambda T) - 1]}$$

$$c_1 = 3,74 * 10^{-16} \text{ W*m}^2$$

$$c_2 = 1,44 * 10^{-2} \text{ K*m}$$

$$\lambda_{\text{max}} * T = 2896 \mu\text{m*K}$$



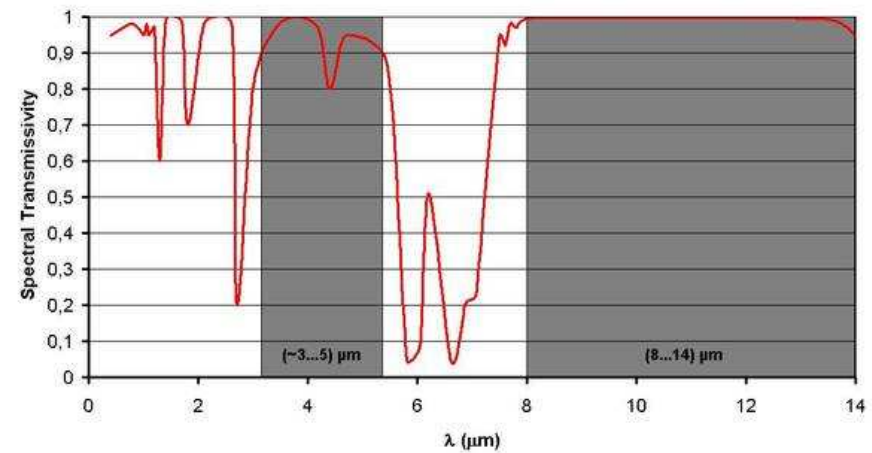
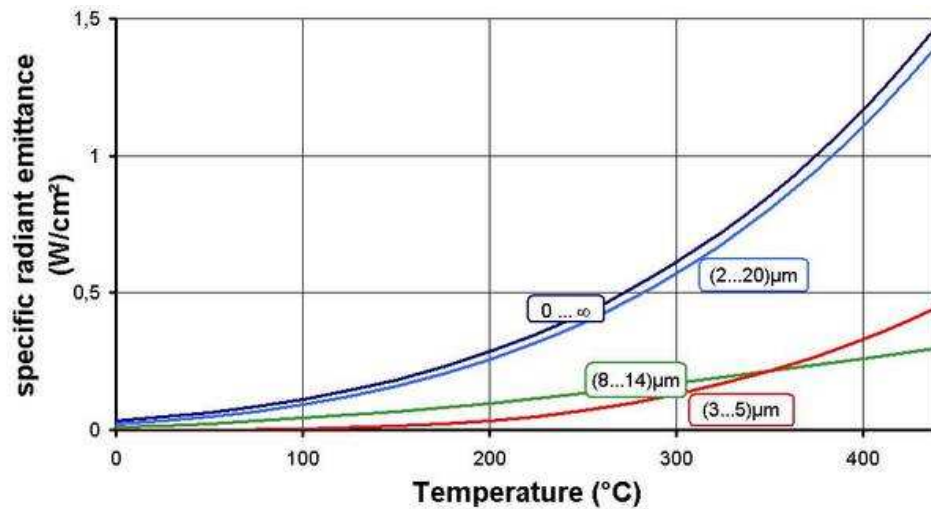
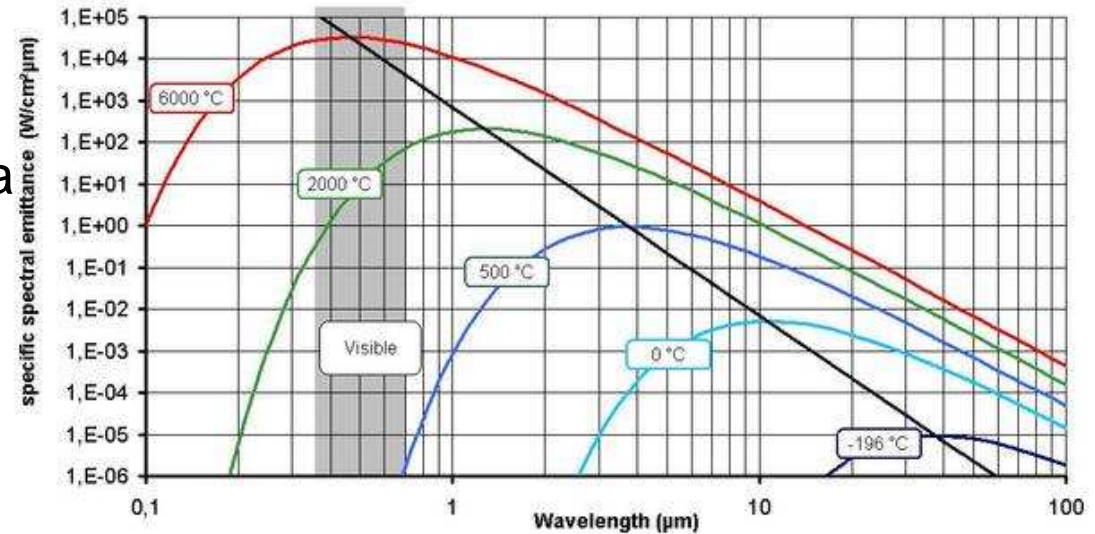


# EM laws TIR

$$E = \epsilon \sigma T^4$$

$\sigma$  - Stefan-Boltzmann constant

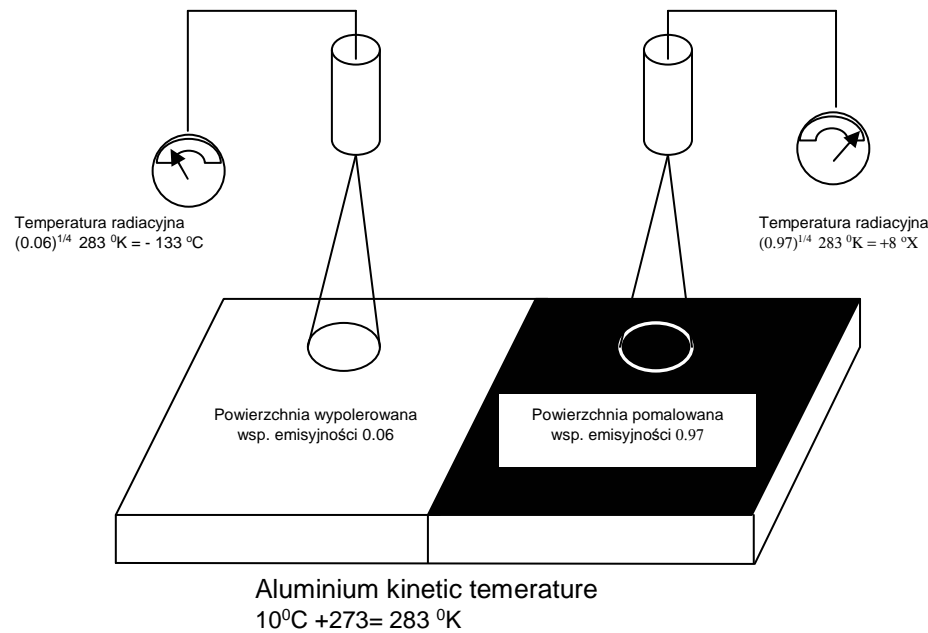
$$5,67 \cdot 10^{-8} \text{ W } / (\text{m}^2 \cdot \text{K})$$



# Map algebra - temperature

## Radiant temperature

$$T = \sqrt[4]{\varepsilon} \cdot T_{rzecz}$$



$$E = \sigma T^4 = \sigma [(\varepsilon)^{1/4} T_{rzecz}]^4 = \sigma \varepsilon T_{rzecz}^4$$

# Map algebra - temperature

## Temperature

$$L_{\lambda} = gain \cdot DN + offset = \left( \frac{L_{\max} - L_{\min}}{255} \right) \cdot DN + L_{\min}$$

$(W \cdot m^{-2} \cdot sr^{-1} \cdot \mu m^{-1})$

- where:
- $L_{\lambda}$  – spectral radiance of the sensor,
- $L_{\min}$  – minimum spectral radiance , in thermal band 0.00,
- $L_{\max}$  – maximum spectral radiance , in thermal band 17.04
- DN – of thermal band

# Map algebra - temperature

$$M_\lambda = \frac{c_1}{\lambda^5 * [\exp(c_2/\lambda T) - 1]}$$

$$c_1 = 3,74 * 10^{-16} \text{ W*m}^2$$

$$c_2 = 1,44 * 10^{-2} \text{ K*m}$$

$$T = \frac{K_2}{\ln\left(\frac{K_1}{L_\lambda} + 1\right)}$$

gdzie:

$L_\lambda$  - luminancja spektralna zarejestrowana przez radiometr satelity  $[W \cdot m^{-2} \cdot sr^{-1} \cdot \mu m^{-1}]$

$K_1, K_2$  - stałe kalibracyjne

$$K_1 = \frac{2 \cdot \pi \cdot c^2 \cdot h}{\lambda^5} = 666.09 \left( \frac{W}{m^2 \cdot sr \cdot \mu m} \right)$$

$$K_2 = \frac{h \cdot c}{k \cdot \lambda} = 1282.71 \text{ (K)}$$

$k$  - stała Boltzmann  $1,380 \cdot 10^{-23} \left( \frac{J}{K} \right)$

$h$  - stała Plancka  $6,626 \cdot 10^{-34} \text{ (J} \cdot \text{s)}$

$c$  - prędkość światła  $2,998 \cdot 10^8 \left( \frac{m}{s} \right)$

$\lambda$  - długość fali (m)

